

# High Redshift Starbursts in the Hubble Deep Field

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**Evolution of Starbursts Conference**

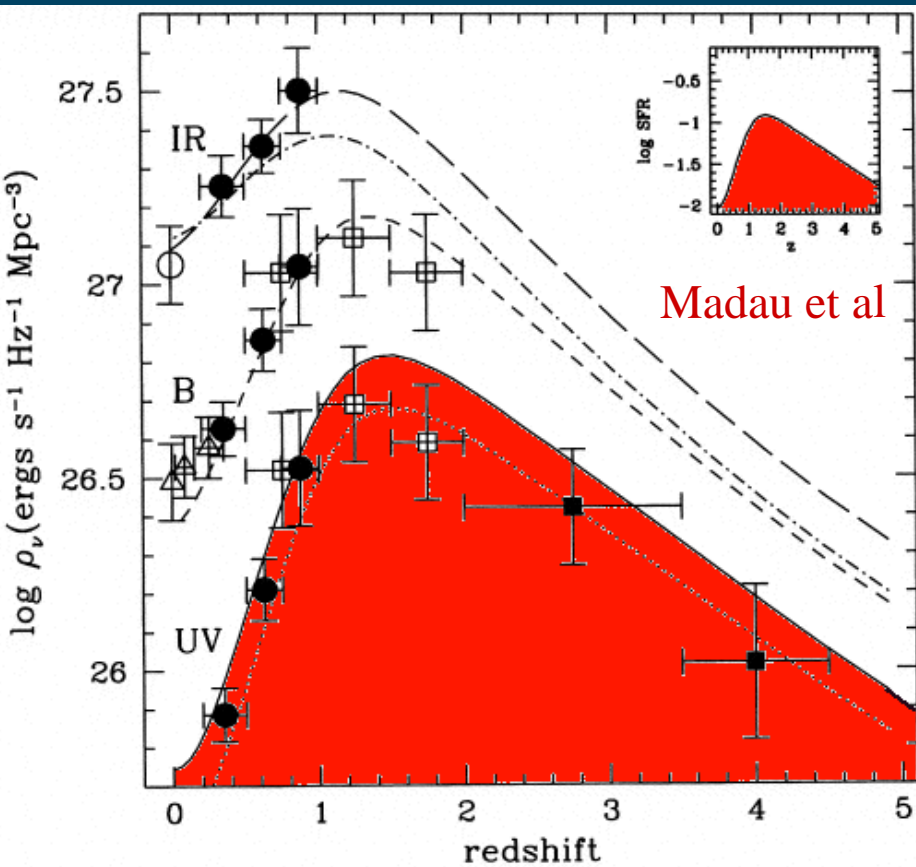
**16<sup>th</sup> – 20<sup>th</sup> August 2004**

**Bad Honnef**

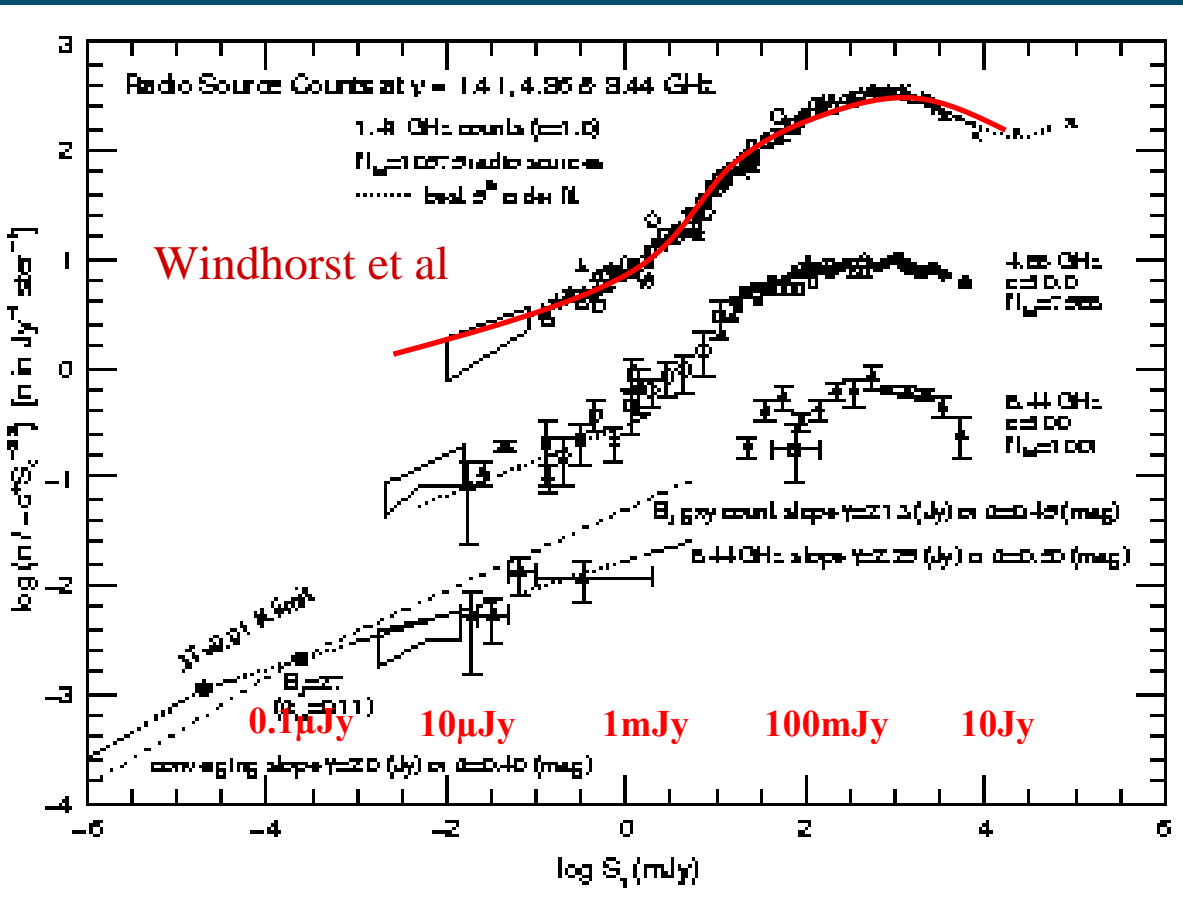


# Star-formation At High Redshift

Infra-red, sub-mm, and radio observations are revealing enhanced star-formation at high redshift. Some of these systems may be dust obscured



# $\mu\text{Jy}$ Radio Source Population



- Differential source counts show an upturn below flux densities of a few mJy
- A new population of starburst systems ?

# The Hubble Deep Field – A Window On The Distant Universe



**Galaxy counts and colours imply that many of these distant galaxies are merging**

- **~2500 distant galaxies in an area covering just 5.3 square arcminutes.**
- **Are we seeing a population of distant starburst galaxies ?**
- **With such a crowded field, astrometric alignment between wavebands is essential in order to obtain reliable identifications**

# Hubble Deep & Flanking Fields



- **HST WFPC2**  
(Williams et al 96)
- **Central HDF**  
– 150 orbits  
complete to  $R > 29$
- **Surrounding HFF**  
– 10 orbits  
Complete to  $R > 25$

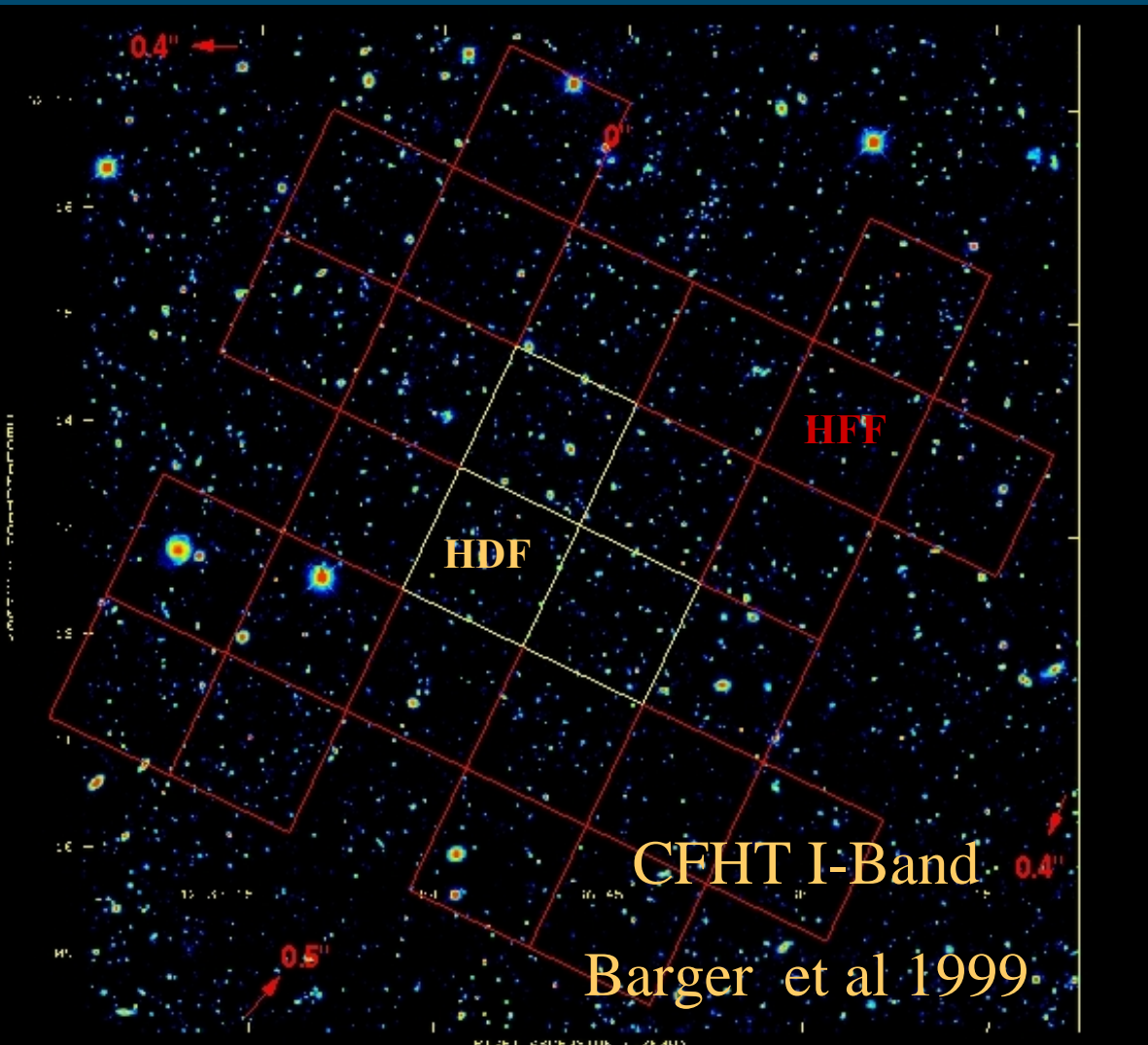
**Hubble Deep Field**

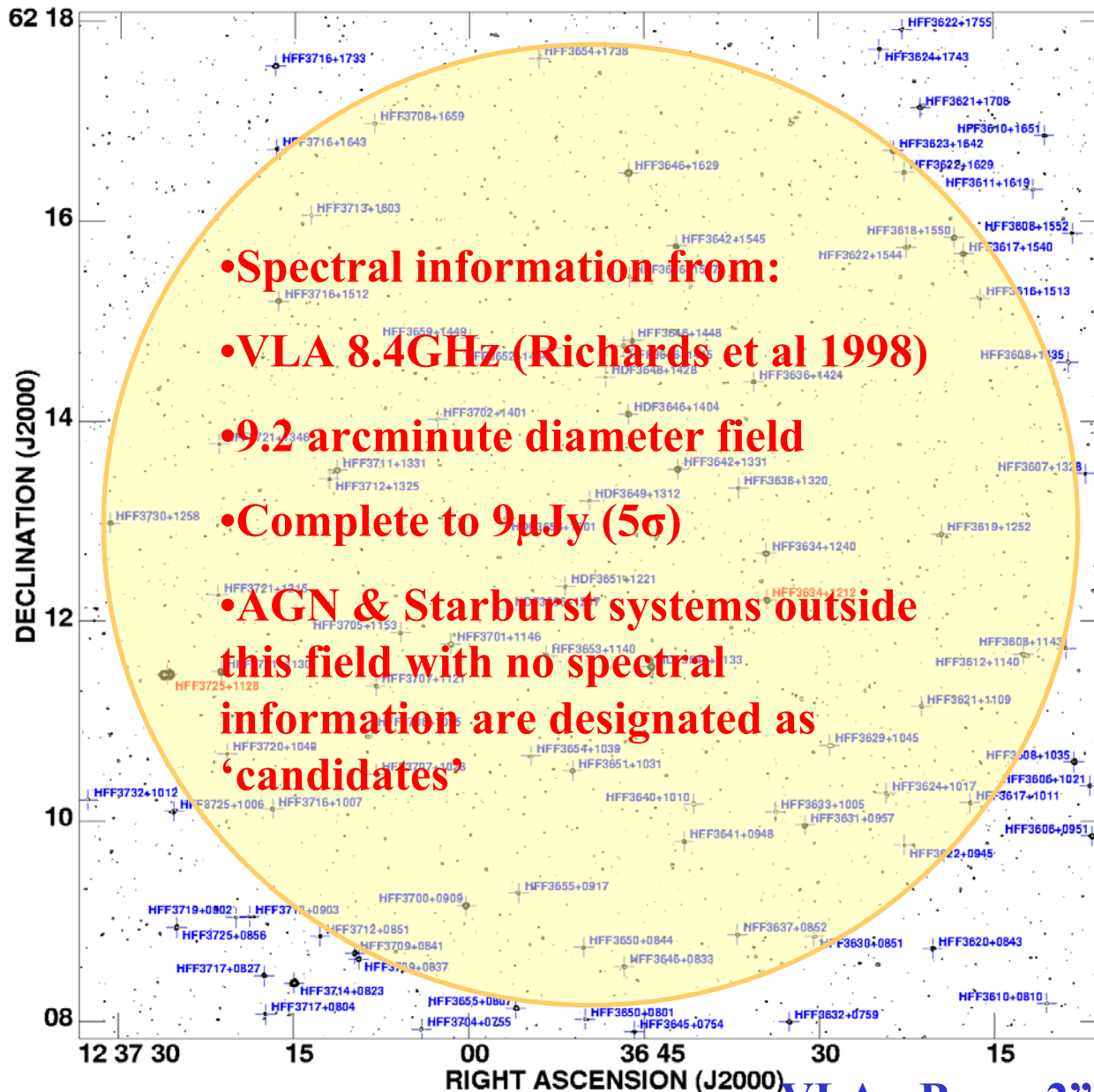
**HST · WFPC2**

PRC96-01a · ST ScI OPO · January 15, 1996 · R. Williams (ST ScI), NASA

# Aligning the HST WFPC2 Images to the Radio Reference Frame (ICRF)

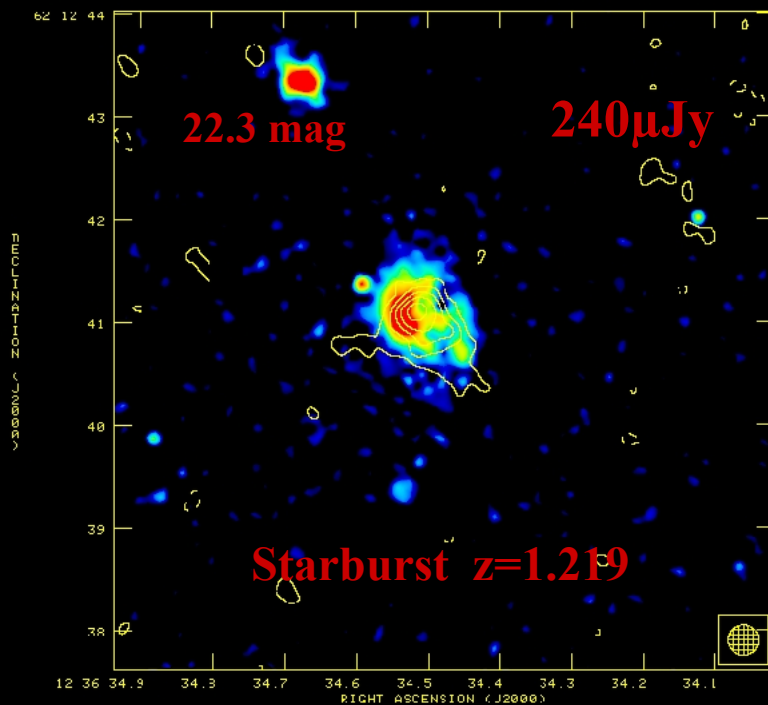
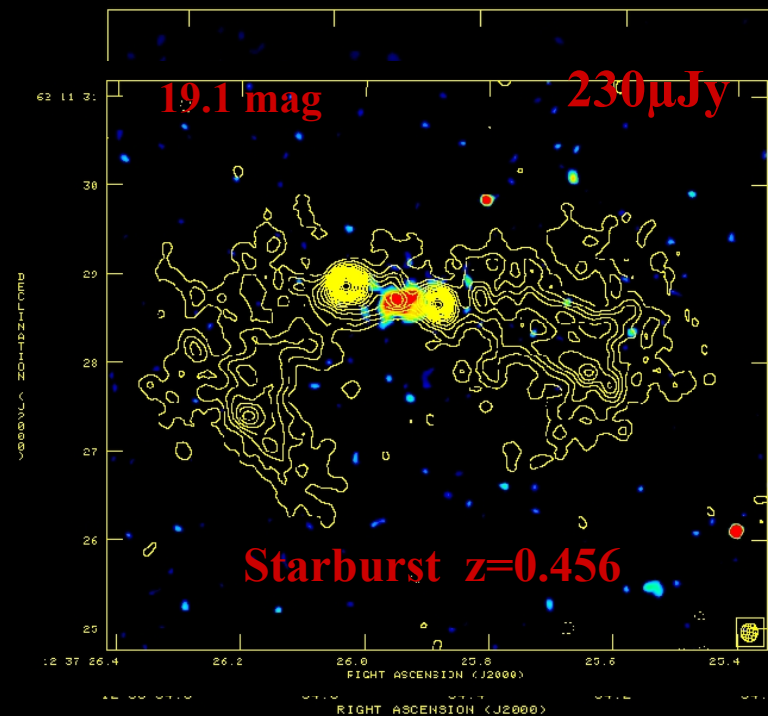
- The HDF and HFF are enclosed in an area of  $\sim 10$  arcminutes square.
- 10 arcminute field mapped by MERLIN+VLA
- CFHT I-Band image (9 arcminute field) aligned under radio sources (4 parameter solution to distortion).
- HDF/HFF frames aligned under CFHT image for compact galaxies.
- Alignment to better than 50mas in HDF, rising to  $\sim 150$ mas in outer HFF & 150-250 beyond the HFF





- Spectral information from:
- VLA 8.4GHz (Richards et al 1998)
- 9.2 arcminute diameter field
- Complete to  $9\mu\text{Jy}$  ( $5\sigma$ )
- AGN & Starburst systems outside this field with no spectral information are designated as 'candidates'

- MERLIN (18 days) + VLA (42 hours) 1.4 GHz image is amongst the most sensitive ever made ( $\sigma \sim 3.2\mu\text{Jy}$ ) Beams 0.2 - 0.5"
- In 10 arcmin FOV there are 92 sources  $> 40\mu\text{Jy}$
- AGN have compact cores and inverted radio spectra
- Starburst systems have steep radio spectra, are extended on sub-galactic scales, and often have I-R detections



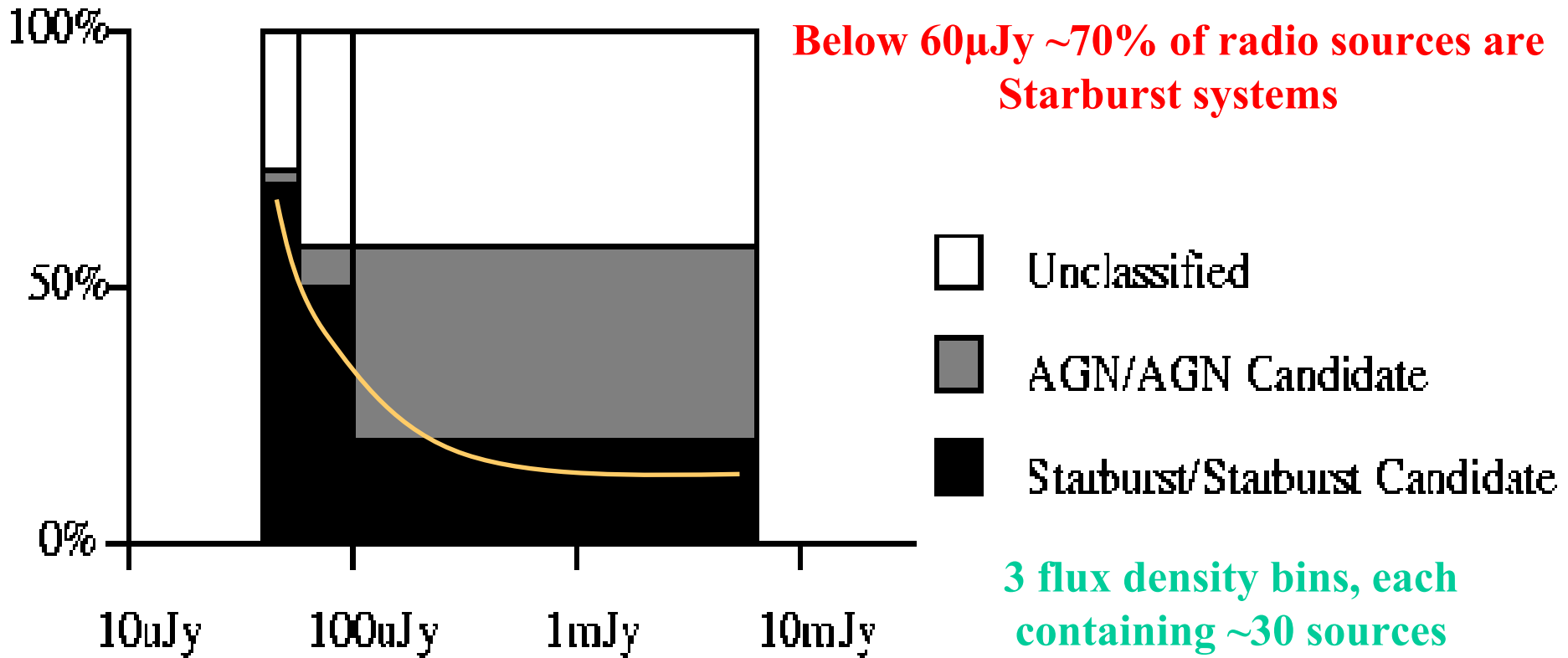
# Radio Structures

- Most have angular sizes in the range 0.2 – 3 arcseconds
- Tail of mJy population show high luminosity radio structures
- ~85% are identified with galaxies brighter than  $R=25$ .
- Most AGN systems show small core-jets (flat-spectrum cores)
- Starbursts have steep radio spectra and are extended
- Sources which cannot be identified as AGN or Starburst are listed as 'Unclassified'

The 15% unidentified sources appear to be very distant dust-obscured starburst systems...

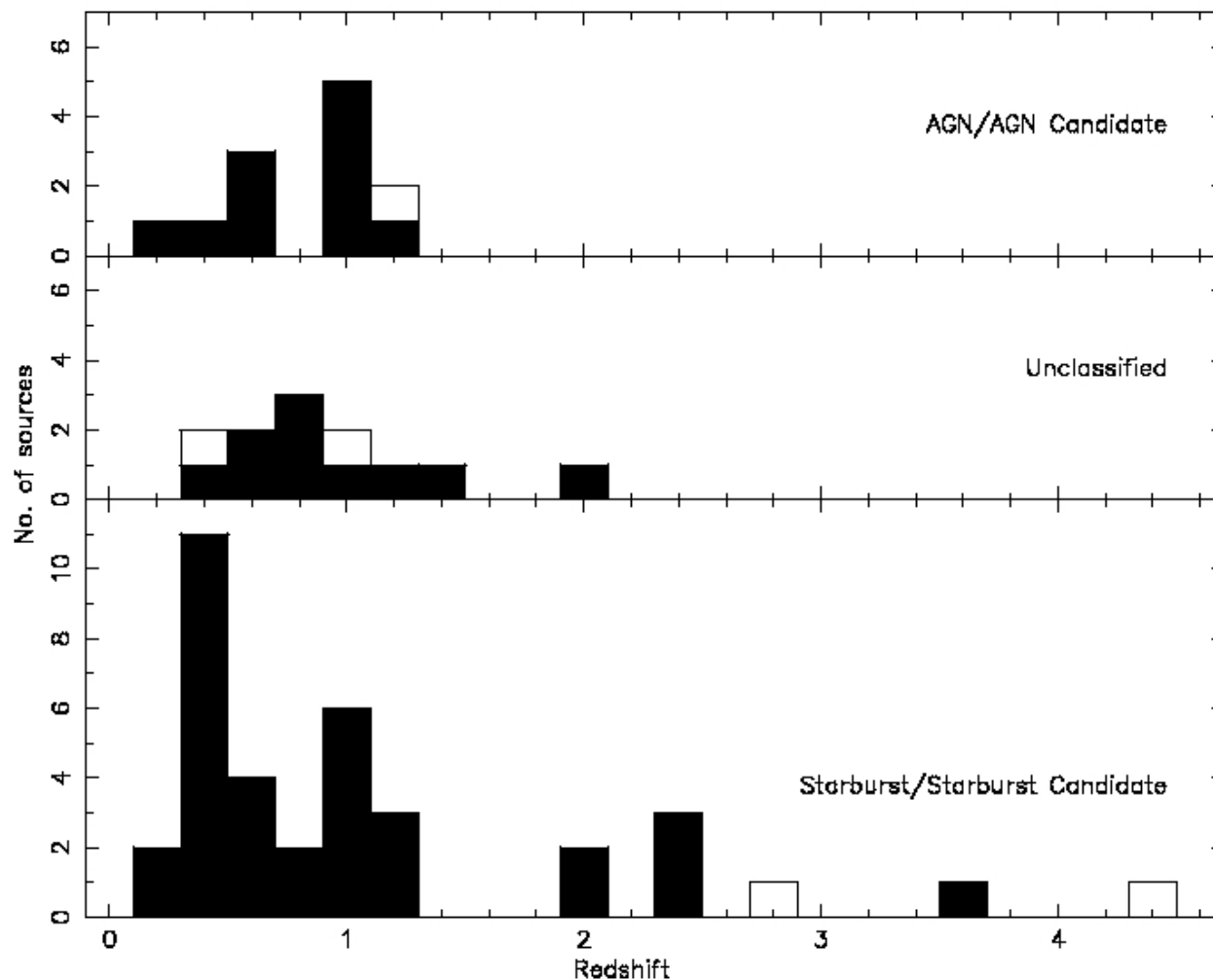


# Distribution of Source Structure with Flux Density



**Increasing proportion of Starburst/Starburst Candidates with decreasing source strength**

# Redshift Distribution



**60 / 92 radio sources in the 10-arcminute field have measured spectroscopic (Black) or photometric (white) redshifts**

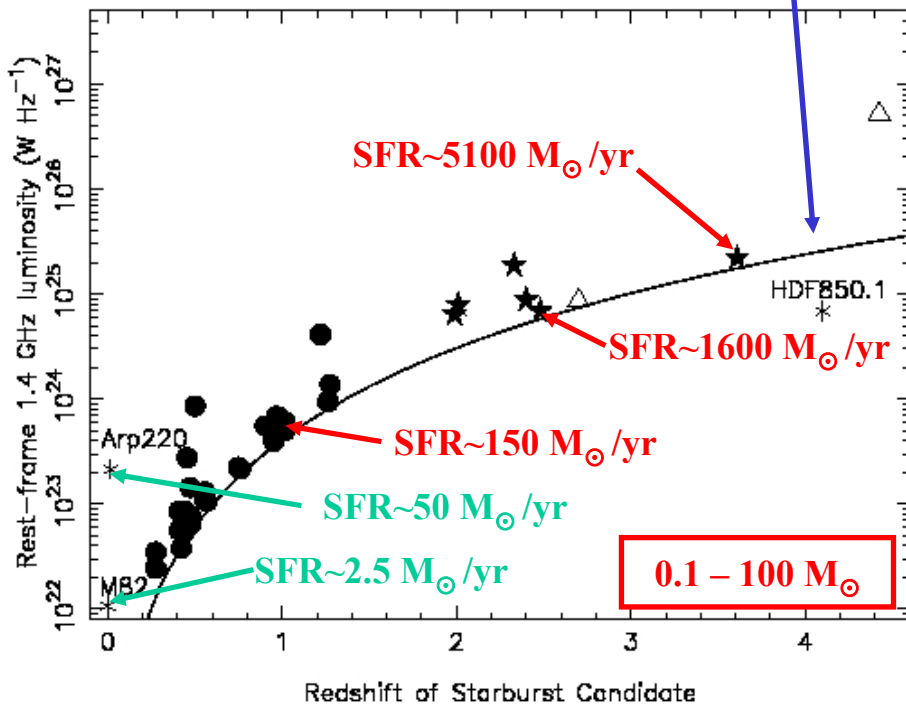
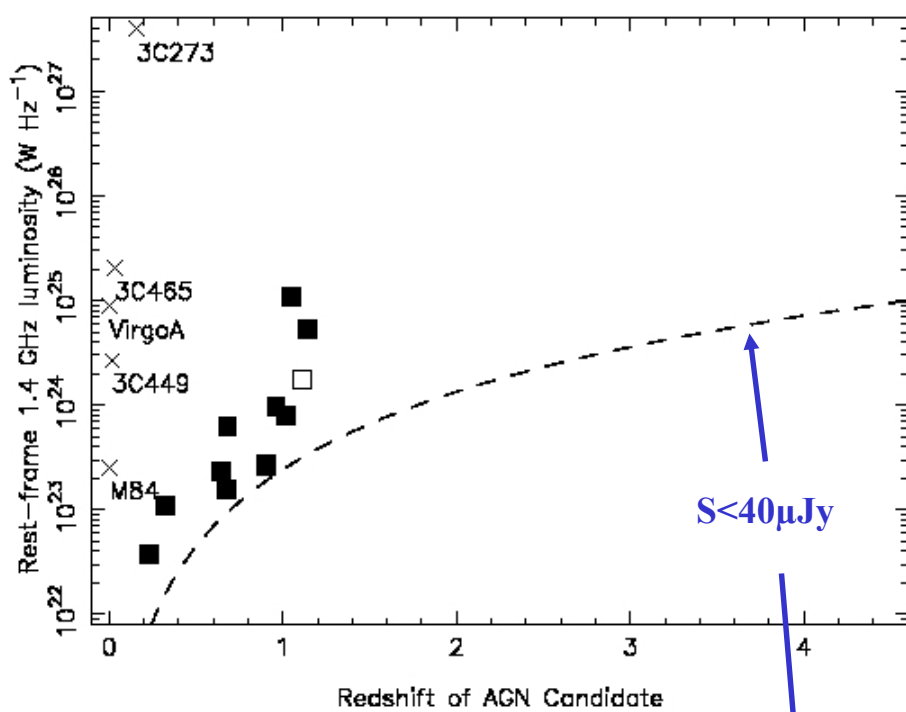
**• Similar redshift distribution for AGN & Unclassified source types – median redshifts of 0.91 & 0.85 respectively**

**• Starburst systems are mostly at lower redshifts ( *nb 2 clusters within the field with  $z \sim 0.4$*  ) with a tail to redshifts  $>4$  !! Median redshift=0.56**

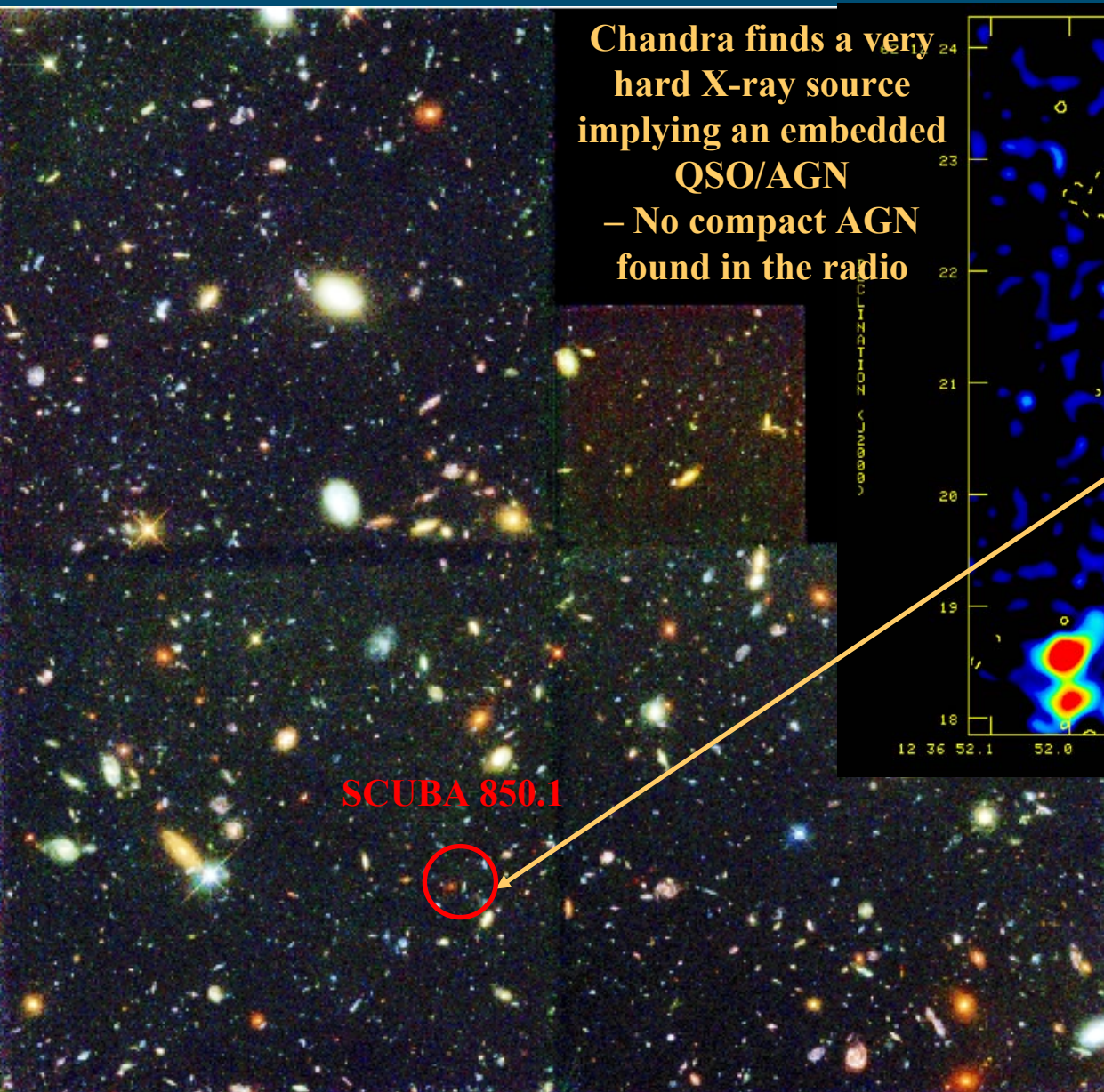
For those sources with measured redshifts (65%):

AGN systems have luminosities typical of lower luminosity 'classical' double AGN systems - although very few are doubles, most are small core-jet sources

Starburst systems have luminosities far greater than M82. Most high redshift starbursts and the sub-mm starbursts (\*) have luminosities orders of magnitude larger than M82 - and even the luminous local starburst Arp220. 2 optically faint EROs ( $\Delta$ ) have very high luminosities and may contain embedded AGN

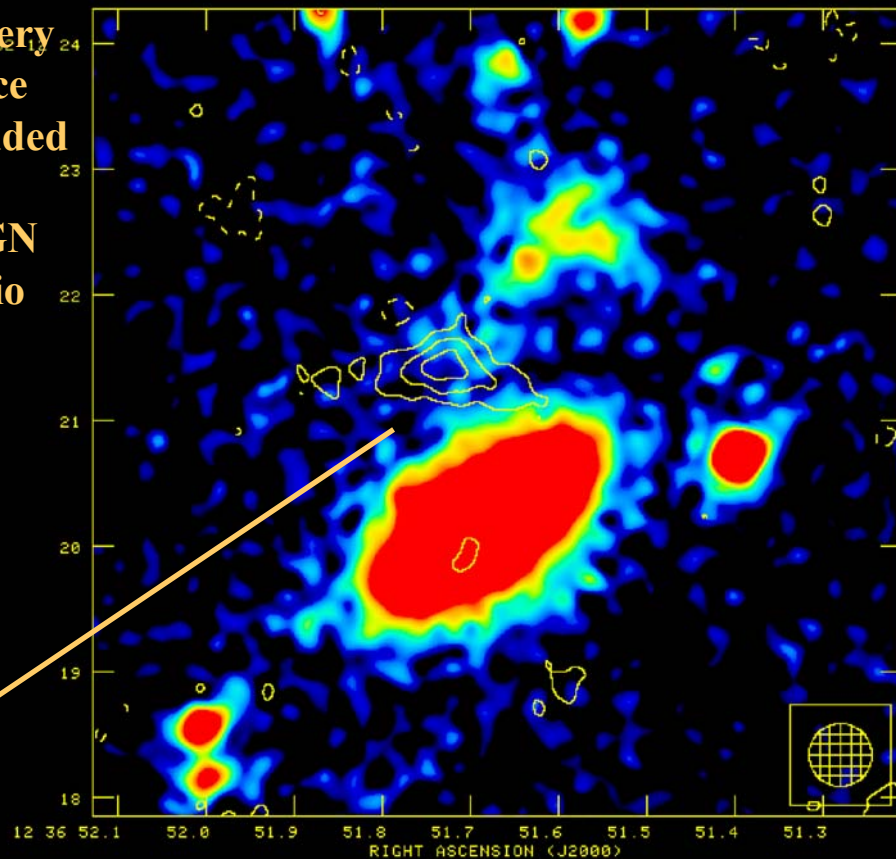


**Starformation at high redshift: – The optically faint systems -- A new population of very high redshift, dust-shrouded starburst galaxies**



**SCUBA 850.1**

**Chandra finds a very hard X-ray source implying an embedded QSO/AGN – No compact AGN found in the radio**



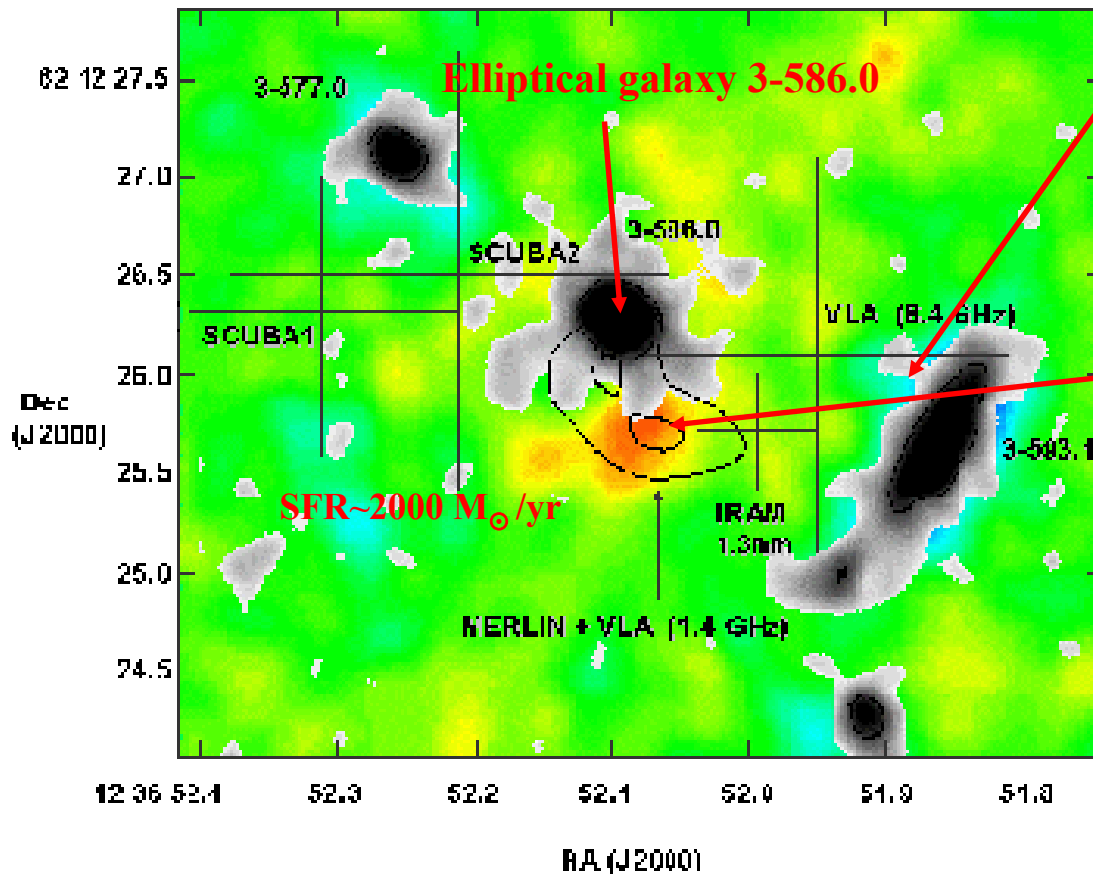
**J123651+621221 at 1.4GHz with MERLIN+VLA – Red NICMOS detection implies  $z \sim 2.6-2.7$  ( $S_{1.4} = 49 \mu\text{Jy}$ ) – (Dickinson et al 2000)**

## Starformation at high redshift: HDF850.1

Subaru K'-band residual image (false colour) after subtraction of WFPC2 image (greyscale) smoothed to the angular resolution of the Subaru image

Also seen by NICMOS in 1.4 $\mu\text{m}$  – 1.8 $\mu\text{m}$  band

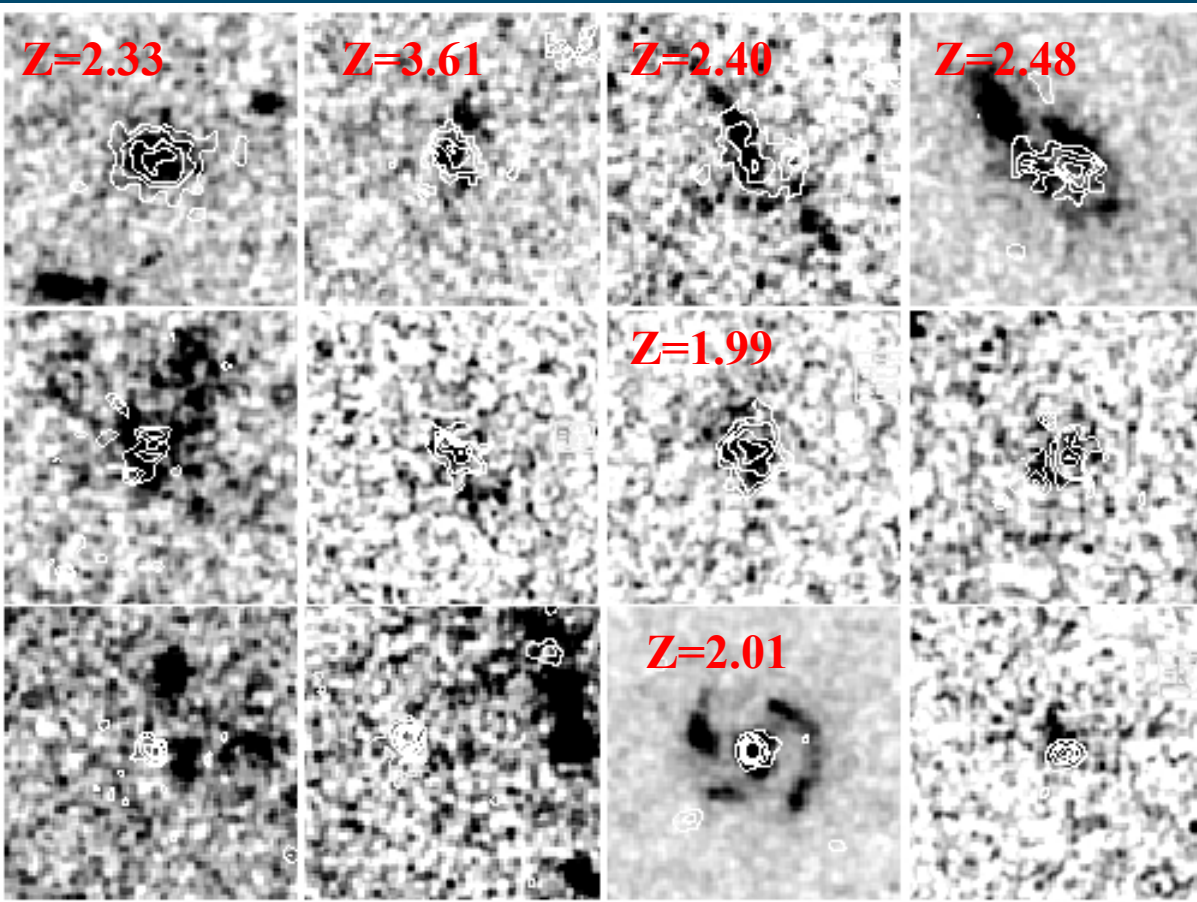
The host galaxy of



- MERLIN(+VLA) observations have allowed the WFPC2 HDF frames to be aligned to the ICRF to better than 50mas
- The IRAM detection of the original SCUBA HDF850.1 sub-millimetre source can now be seen to be associated with an optically faint system lying beyond the limit of the HST HDF observations
- Radio emission close to the detection limit with present sensitivity ( $4\sigma \sim 16\mu\text{Jy}$ ) has been found from close to this position together with a very red object suggesting a distant starburst system at a redshift  $\sim 4.1$  – derived from the shape of the SED

# Sub-mm / Radio Source Associations

Chapman et al (2004) have identified a number of brighter sub-mm sources (in a field centred in the HFF to the North of the HDF) with radio sources in the 10-arcminute field



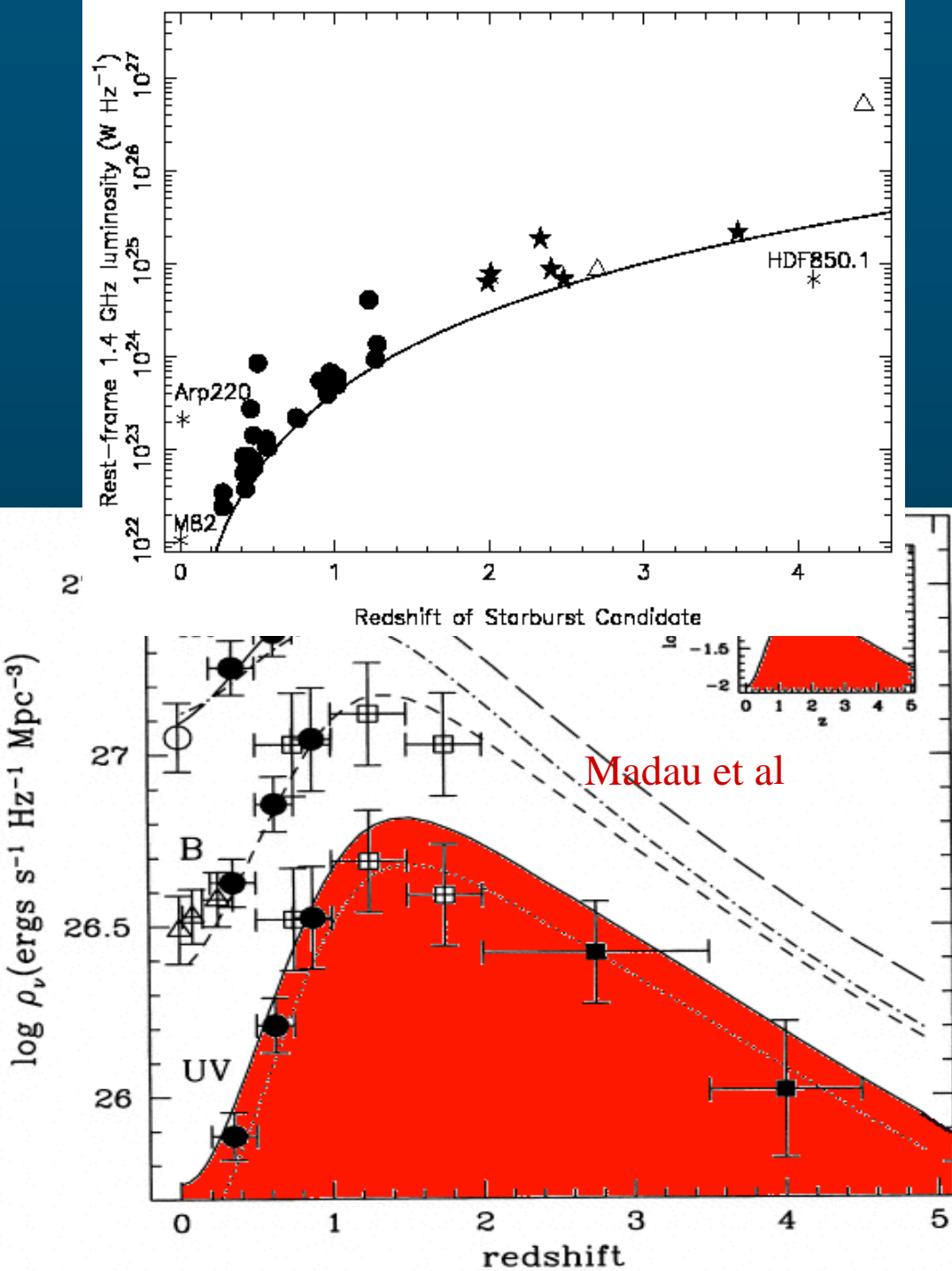
Optical counterparts are typically fainter than most Starburst systems in the 10-arcminute field ( $22 < I_{\text{mag}} < 26$ ) and lie at significantly higher redshifts

Redshifts shown for those sources with  $S_{1.4\text{GHz}} > 40\mu\text{Jy}$

*Contours MERLIN+VLA 1.4GHz, greyscale HST STIS/ACS*

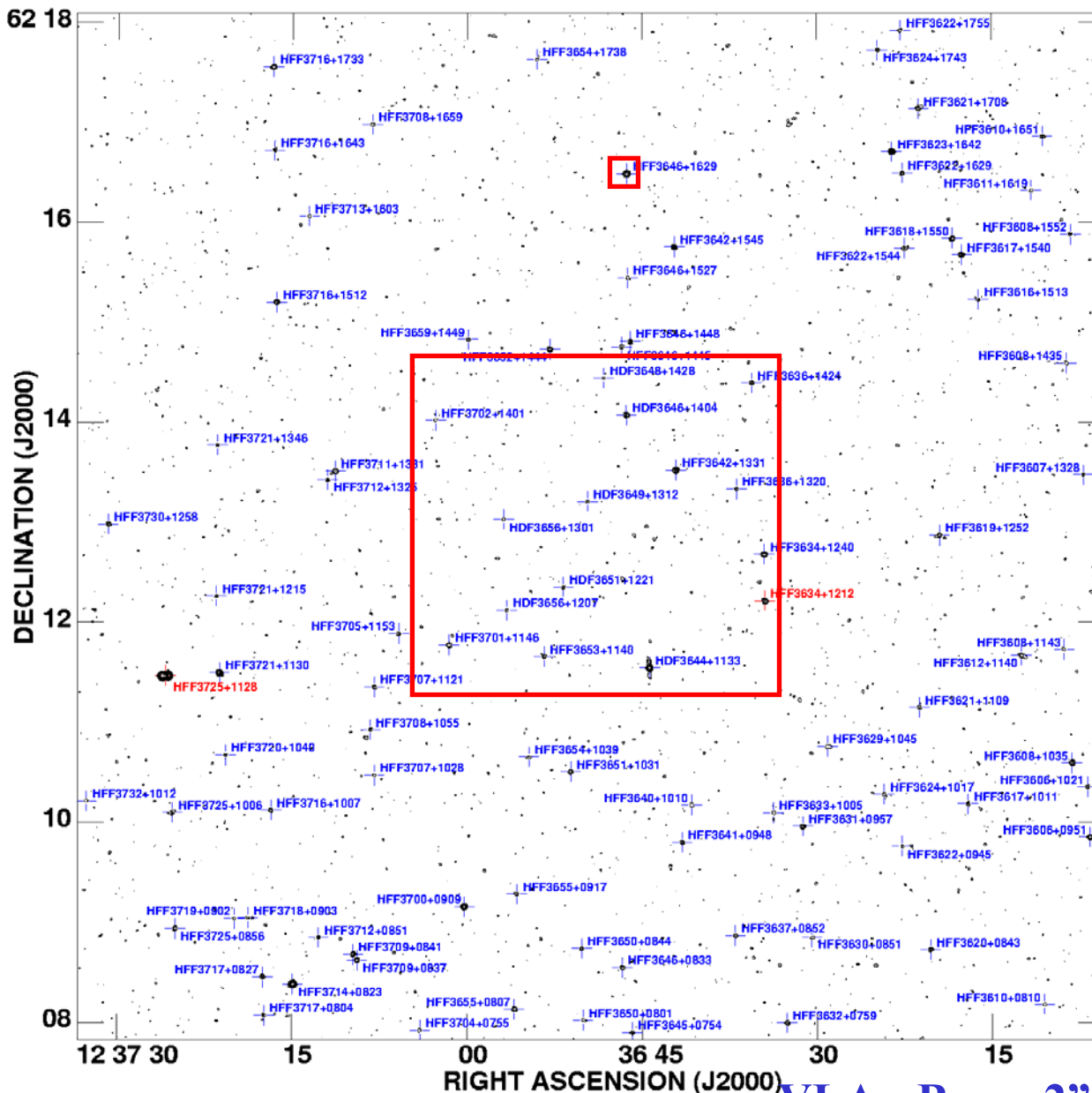
# Star-formation At High Redshift

- Most Starburst systems in the 10-arcmin field (●) are galaxies at redshifts less than  $z=1.5$
- The sub-mm Starbursts (\*) are at high redshifts typically in excess of  $z=2$
- The optically faint, dust obscured systems (▲) also lie at high redshift. Some of these may be Starbursts with embedded AGN
- We are only beginning to sample the population of high redshift star-forming galaxies



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# The 3 Arcminute Field



- Central 3.4x3.8 arcminutes of the field have been fully mapped with combination MERLIN+VLA images
- Statistical analysis of radio/optical positions made

VLA - Beam 2''

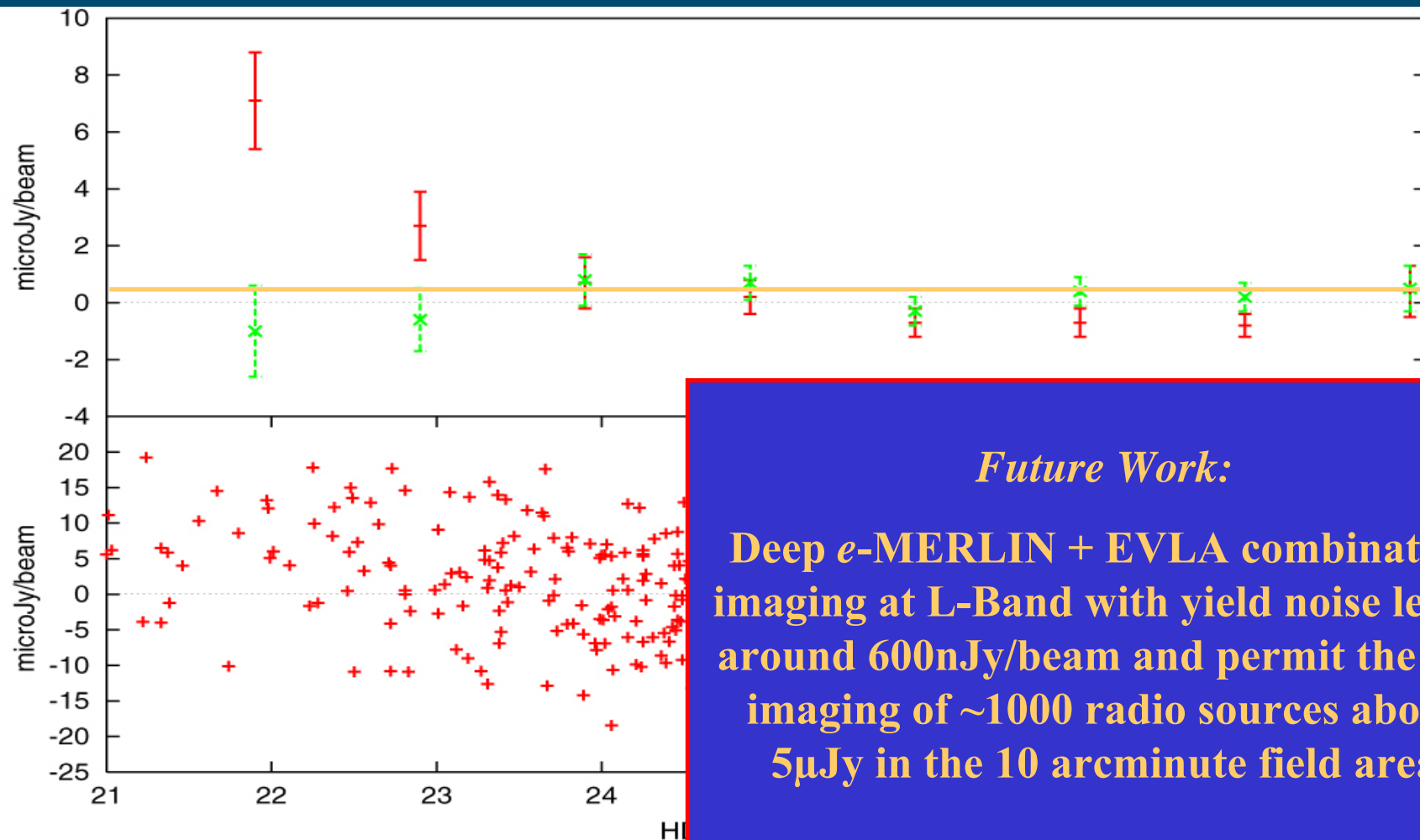


# Radio brightness at 1.4GHz (smoothed to 1-arcsec resolution) at the positions of known galaxies plotted against WFPC2 I-magnitude

Control data incorporate a random 7-arcsec shift applied to the radio position

Upper panel – binned by I-magnitude

Lower panel – data shown unbinned



*Future Work:*

Deep *e*-MERLIN + EVLA combination imaging at L-Band with yield noise levels around 600nJy/beam and permit the full imaging of  $\sim 1000$  radio sources above  $5\mu\text{Jy}$  in the 10 arcminute field area

3-Arcminute Field – Statistical Analysis of Radio/Optical Positions

# Conclusions:

- Radio data have been used to astrometrically align multi-waveband images of the HDF and HFF allowing unambiguous identifications for the first time
- Most radio structures have angular sizes in the range 0.2 to 3 arcseconds
- AGN systems are nearly all small core-jet structures
- Proportion of starbursts increases with decreasing flux density – below  $60\mu\text{Jy}$  >70% of sources are starburst type systems
- ~85% of radio sources identified with galaxies brighter than  $I=25\text{mag}$
- ~15% optically faint systems are thought to be mostly dust obscured high redshift starbursts, some of which may contain embedded AGN
- Most starburst galaxies are associated with galaxies at redshifts less than  $z=1.5$
- Sub-mm source identifications bedevilled by poor sub-mm astrometry
- Identified sub-mm starburst systems are associated with galaxies which are fainter than most HDF/HFF starbursts and found at high redshift with  $z$  typically  $> 2$

