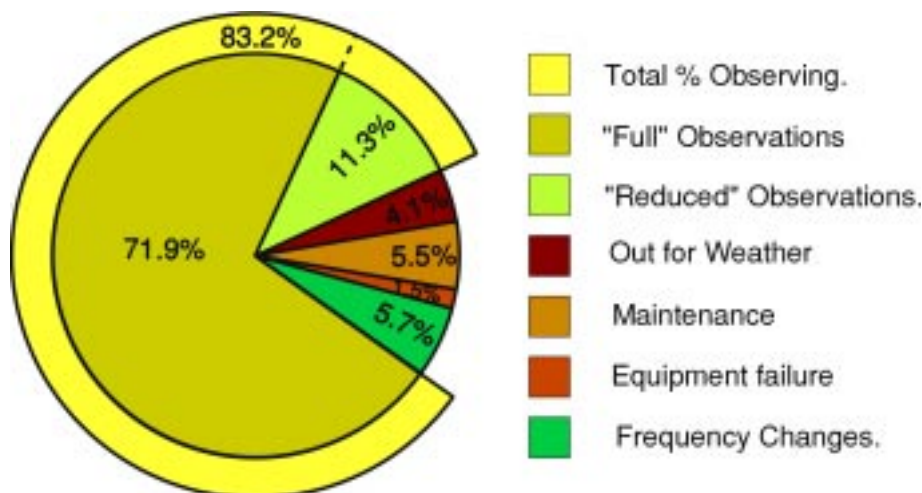


The MERLIN target for operation is normally 24 hours per day for 9 months of the year except for a weekly maintenance period not exceeding 8 hours and the time required for PATT-requested frequency changes. During the remaining 3 months of the year, usually in the summer/early autumn, major mechanical maintenance work, painting of telescopes and development takes place. A frequency change also usually takes place during this 3 month period. During the two years covered by this report, the MERLIN Steering Committee authorised extensions of the summer maintenance period to four months, to facilitate major engineering works as part of the restructuring programme. Efficient scheduling of the engineering work held the actual down time to 3.8 months, 0.75 months of which was used for EVN observations by the Mk2 and Lovell telescopes. The frequency flexibility implemented as part of this engineering programme is already paying dividends when considering the number of K-Band astronomical programmes that are being completed. The new ability to switch almost instantaneously between C-Band and K-Band has resulted in more K-Band proposals being completed during the past 2 years than ever before. When one considers the appalling weather during these 2 years, this represents a major improvement in the operational efficiency of MERLIN. In the past it could have taken a few weeks to make this frequency change with no guarantee of good weather. With further investment, the benefits of frequency flexibility can be extended to include L-Band. This is already available on the Cambridge and Mk2 telescopes and will be implemented at Defford from Autumn 2001.

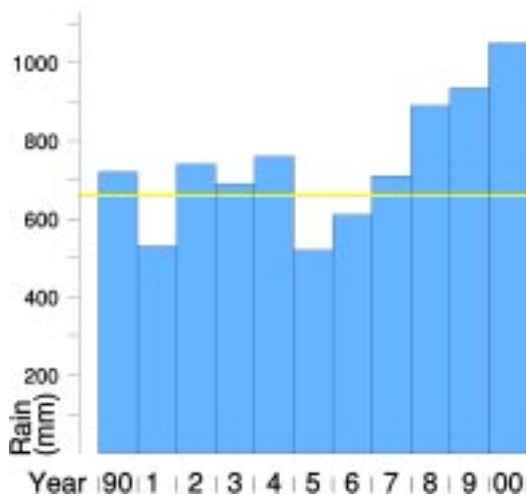
Operating Period

A summary of the operational status of MERLIN is given in the pie chart for the period of time which was available for observations during the Calendar years 1999 and 2000. Included in the time for maintenance is almost a day at the millennium, during which time MERLIN was not operated for safety reasons in case of possible computer or utility failure. As a result of a systematic and comprehensive programme to identify and correct potential Y2K faults, none occurred within the MERLIN system. The sector labelled 'reduced' observations indicates the sum total of short



Statistics

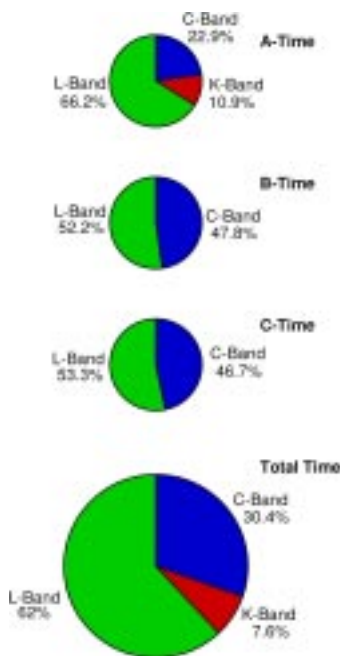
Left: Summary of the operational status of MERLIN during 1999 and 2000.



Above: Rainfall figures for the Jodrell Bank site from 1990 to 2000.

periods when some fault or bad weather affected or prevented the operation of a single telescope, the loss of which was not considered to have a significant effect on the final images produced other than a degradation in the signal-to-noise. This figure is significantly higher than in previous periods primarily due to 'teething' problems with the new drive system of the Darnhall telescope - the first to be replaced. Low-level faults continued to occur over many months before their cause was finally eradicated by the National Facility engineers in conjunction with the company supplying the drive system. It can be seen that, including this time, observations were made for 83.2% of the time. The 4.1% down time caused by weather is, as for the period of the previous biennial report, once again unusually high in comparison with that prior to the past four years. The highest ever recorded rainfall in 1998 was superseded in 1999, which in turn was further superseded in 2000, during which year it was over 400 mm above the yearly average at the Jodrell Bank site. Although rain does not prevent the actual operations of the telescopes, it does result in a degradation of observations at the higher frequencies.

During Semesters 99A, 99B, 00A and 00B, MERLIN has operated in its three main frequency bands; L-Band, C-Band and K-Band. The actual time spent in observations within each band has been set by astronomical demand as determined by the MERLIN Time Allocation Group, the EVN Programme Committee for the MERLIN+VLBI observations and, in the case of K-Band observations, the state of the weather. As indicated earlier, the value of the 'frequency flexibility' programme was demonstrated in the number of K-Band proposals completed, more than ever before. The apparent imbalance in the time allocated to the different observing bands has been due to three L-Band programmes classified as 'key-programmes' which required very deep integrations and hence a lot of time, and also to a greater demand for L-Band observations. However, it is already known that for the year 2001, C-Band proposals will dominate.



Above: Summary of MERLIN time allocations for each observing band and proposal rating.

Of the programmes allocated time by PATT during 1999-2000, 92% of the A-priority and 86% of the B-priority were completed (Appendix B). Less than 100% completion of the L-Band programmes allocated time was primarily due to interference preventing observations at some frequencies corresponding to redshifted hydrogen. Interference at L-Band frequencies is increasing all the time, though it has recently been possible to have some of the interfering signals switched off for periods of time. Weather prevented the K-Band A-priority programmes from being completed. The adjacent pie charts show the distribution of actual observed time by observing band.

The flexible scheduling policy for MERLIN means that a number of time-constrained observations, such as those of a new nova, can be made at short notice. The MERLIN Director also has at his disposal 2 days per Semester for observations of an immediate urgency. Such Target-of-Opportunity observations are provisionally allocated by PATT and, when made, are classified as A-priority observations and have been included as A-priority in the pie-charts.